

Apollo 17 ALSEP ARCSAV Heat Flow Experiment Calibrated Gradient Bridge Temperatures Collection (1975-092 to 1975-181)

The Apollo 17 astronauts deployed two heat flow probes as part of the Apollo Lunar Surface Experiment Package (ALSEP) [1a]. These probes operated from December 1972 to September 1977 [1b]. Data obtained from December 1972 to December 1974 were processed by the original investigators of the Heat Flow Experiment (HFE) [2] and are available through PDS [3]. This collection contains the Apollo 17 HFE data for April through June 1975. They were extracted from the original ALSEP raw data archival tapes (also known as 'ARCSAV tapes') [4]. These tapes were recovered in 2010 at the Washington National Records Center [5]. The HFE raw data packets as extracted from these tapes are available as Level-1 raw products [6].

This collection contains fully processed, calibrated data for the temperature sensors that were collectively called 'gradient bridges'. The nomenclature of the sensors follows that by the original HFE investigators [2, 7]. The same information is also reviewed by [8]. Each gradient bridge included a pair of resistance temperature detectors (RTD), placed at two different depths into the hole where the heat flow probe was inserted. The heat flow instrument obtained the average and the temperature difference between the paired RTDs of each gradient bridge. Two modes of measurements were used for the temperature difference: low gain and high gain. The low gain mode had a range of ~20 K, and the high gain mode had a range of ~2 K.

For these data products, the temperature values for each RTD have been obtained from the average temperature and the high-gain temperature difference measurements for each gradient bridge RTD pair. When the temperature difference exceeded the range of the high gain mode, the low gain mode was used for obtaining the temperature of each RTD.

These temperature values have been calculated according to the data reduction procedure used at the Johnson Space Center in the 1970s, described in [9] and [10].

The RTD calibration data are also given in [10]. Seven sets of heat flow probes (SN-1 through -7) were fabricated in preparation for the Apollo Heat Flow Experiments, and they were calibrated in 1968 – 1971. Because the calibration work predated the Apollo missions, the calibration reports for these probe sets [10] only refer to their serial numbers assigned by the contractors who fabricated them. These reports do not refer to the actual Apollo flight missions the individual probe sets were on. According to notes left behind by the HFE investigators and the contractors, the serial number for each pair of heat flow probes match up with each Apollo flight as follows.

Apollo Flight	HFE serial number	HFE Flight Model number
Apollo 13	SN-5	2S
Apollo 15	SN-4	2
Apollo 16	SN-6	3
Apollo 17	SN-7	4

The original HFE investigators, based at Columbia University, used a data processing scheme slightly different from the one used at JSC. It is outlined in [11] and [12]. However, these reports did not

provide sufficient information for fully reconstructing their data processing procedure for the Apollo 17 heat flow probes. The data provided here were processed by the JSC scheme.

HFE operational status is reported in the weekly logs of the ALSEP instruments [13].

Data Organization

Each of the two heat flow probes deployed at the Apollo 17 site had two gradient bridges, 'TG11' and 'TG12' for Probe 1, and 'TG21' and 'TG22' for Probe 2 (refer to [2] and [5] for the nomenclature; [3] uses the nomenclature 'DTG11' and 'DTG12' for Probe 1, and DTG21' and 'DTG22' for Probe 2). 'TG11' and 'TG21' are the upper gradient bridges, those on each probe closest to the lunar surface. There are two files in this data package. One file contains data from TG11, and the other contains data from TG21. Data for the lower gradient bridges TG12 and TG22 were not processed, because the use of the calibration data reported in [10] did not yield temperature values that are consistent with the values reported in [3], while those for TG11 and TG21 did. See [8] for more detail.

The two files have an identical, fixed-width, ASCII tabular data organization. In each file, data are organized in six columns. The first column indicates the earth received time, which is typically used as the time of acquisition for analysis purposes. The second column gives the temperature value of the upper RTD of the gradient bridge. The third column gives the temperature value of the lower RTD. The fourth column gives the average of the two. The fifth column gives the temperature difference between the two RTDs obtained in the high-gain mode. The sixth column gives the temperature difference in the low-gain mode. Each row contains the temperature values acquired in one data sampling cycle. All the temperature values are in Kelvin, rounded to the nearest 0.001 K which is the accuracy of the differential temperature measurement [7].

For example, the file for gradient bridge TG11, a17_hfe_1975_l2_arcsav_tg11.tab, is organized in the following manner:

time	TG11A	TG11B	TG11avg	DTH11	DTL11
1975-092T00:04:00.817	256.075	253.158	254.617	2.620	2.917
1975-092T00:11:15.509	256.075	253.158	254.617	2.620	2.917
1975-092T00:18:30.201	256.075	253.158	254.617	2.620	2.918

Data Users' Notes

- Although all raw thermocouple measurements, numbered DH-14/24/34/44 for probe 1 and DH-16/26/36/46 for probe 2 [4], were extracted from the ARCSAV tapes and archived [6], the reduced temperatures for the thermocouples are not included in this collection because the data providers [8] were still looking for calibration data as of late 2018.

Please note two numbering conventions exist in the literature for the thermocouples. This collection follows the nomenclature of original HFE investigators [2, 7], where the thermocouple at the top of probe 1 is designated as TC11, followed successively in the cable by TC14, TC13, and TC12. Another convention, used in [3], designates the thermocouple at the top of probe 1 as TC14, TC13 is the cable thermocouple closest to probe 1, followed by TC12 and TC11.

- The times in the first column of the data files can be reformatted from YYYY-DDDTHH:MM:SS.sss to decimal day by extracting the day of year (DDD), hour (HH), minute (MM), and seconds (SS.sss) strings from the timestamp, converting them to floating point, then performing this calculation:

decimal day = DDD + (HH*3600. + MM+60. + SS.sss) / 86400.0

where 86400.0 is the number of seconds in a day. Pay careful attention to the number of digits the computer carries. For example, to preserve the millisecond-resolution of the timestamps, 11 decimal digits (or 35 binary digits) need be kept, and therefore the variables DDD, HH, MM, and SS.sss may need to be defined as double-precision floating point. Using single-precision floating-point variables to reformat the timestamps would reduce the resolution to about a second.

References

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