

**DIVINER LUNAR RADIOMETER EXPERIMENT (DLRE)
ALIGNMENT REPORT**

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DLRE Alignment Report

I. Introduction

This document outlines alignment tests and preliminary analysis performed on the DLRE instrument before delivery to LRO. The purpose is to document the tests performed to date and to verify alignment requirements. After alignment of the instrument to the spacecraft, further analysis will determine DLRE pointing with respect to the LRO spacecraft axes as a function of azimuth and elevation position. During the commissioning phase of the mission, scans of the lunar horizon and comparison to LROC and LOLA data will provide additional alignment information that can be used to refine the pointing knowledge.

II. Relevant Documents

LRO Pointing and Alignment Spec GSFC 431-SPEC-000113

DLRE Functional Requirements Document JPL D-32375

DLRE Unit History Log #5 (Theodolite Measurements)

JPL AIDS 303739 (Theodolite Measurements)

DLRE Unit History Log 7 JPL D-43204 (Laser Tracker Measurements)

JPL AIDS 304194 (Laser Tracker Measurements)

III. Requirements to be Verified

DLRE-MISC-NEW	Diviner shall stay within its pointing accuracy allocation as stated in Table 5-1 of the LRO Pointing and Alignment Spec (431-SPEC-000113).
DLRE-MISC-NEW	Diviner shall stay within its pointing knowledge allocation as stated in Table 5-2 of the LRO Pointing and Alignment Spec (431-SPEC-000113).
DLRE-MISC-NEW	Diviner shall stay within its pointing calibration allocation as stated in Table 5-2 of the LRO Pointing and Alignment Spec (431-SPEC-000113).
DLRE-FRD-62	Alignment control requirements for the boresights of telescopes A and B relative to the instrument axes shall be better than ± 2.0 mrad in pitch, yaw, and roll ($\pm 1/3$ rd IFOV of a pixel composed of two neighboring detectors).
DLRE-FRD-63a	The orientation of the azimuth axis relative to the alignment cube shall be measured to an accuracy of ± 0.1 mrad.
DLRE-FRD-63b	The orientation of the elevation axis relative to the alignment cube shall be measured to an accuracy of ± 0.1 mrad.
DLRE-FRD-63c	For a single orientation in azimuth and a nominally 90° orientation in elevation, alignment of the boresight of telescopes A relative to the alignment cube shall be measured to better than ± 0.6 mrad.
DLRE-FRD-63d	For a single orientation in azimuth and a nominally 90° orientation in elevation, alignment of the array axes relative to the instrument axes shall be measured to better than ± 6 mrad. The axis of an array is defined by the best fit line through the peaks of all the detectors in that array.
DLRE-FRD-240	The plane of the solar calibration target shall be inclined 15° above the spacecraft XY plane

Table 1. DLRE alignment requirements.

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Table 5-1. Diviner Pointing Accuracy Allocation

		Arc-Sec			micro-radian		
		Rx	Ry	Rz	Rx	Ry	Rz
ACS	ACS Knowledge	30	30	30	145	145	145
	ACS Controller	15	15	15	73	73	73
Jitter	Jitter	50	50	50	242	242	242
Thermal	Thermal Distortion	200	200	200	969	969	969
Static Bias	1-g Release	150	150	150	727	727	727
	Launch Shift	30	30	30	145	145	145
Static Bias RSS	Ground Alignment Accuracy	500	500	500	2423	2423	2423
		523	523	523	2534	2534	2534
Instrument		413	413	413	2000	2000	2000
Total		1231	1231	1231	5963	5963	5963
Requirement		1238	1238	1238	6000	6000	6000
Unallocated margin		8	8	8	37	37	37

Table 2. DLRE pointing accuracy allocation. Table 5-1 from LRO Pointing and Alignment Spec GSFC 431-SPEC-000113.

Table 5-2. Diviner Pointing Knowledge Allocation

		Arc-Sec			micro-radian		
		Rx	Ry	Rz	Rx	Ry	Rz
ACS	ACS Knowledge	30	30	30	145	145	145
Jitter	Jitter	50	50	50	242	242	242
Thermal	Thermal Distortion	200	200	200	969	969	969
Calibration Error		124	124	124	600	600	600
Instrument		124	124	124	600	600	600
Total		528	528	528	2557	2557	2557
Requirement		619	619	619	3000	3000	3000
Unallocated margin		91	91	91	443	443	443

Table 3. DLRE pointing knowledge allocation. Table 5-2 from LRO Pointing and Alignment Spec GSFC 431-SPEC-000113.

IV. Nominal Configuration

Figure 1 shows schematically the DLRE instrument in its stowed position with relevant axes and vectors. The alignment cube is fixed to the base, so cannot rotate with respect to the spacecraft axes. Nominally the azimuth axis is parallel to the spacecraft Z axis, the elevation axis is perpendicular to the azimuth axis and, when the instrument is stowed, is parallel to the cube face 2 normal vector. The cube face 1 normal vector is perpendicular to both the azimuth axis and to cube face 2. Table 4 lists the nominal unit vectors with respect to the spacecraft axes.

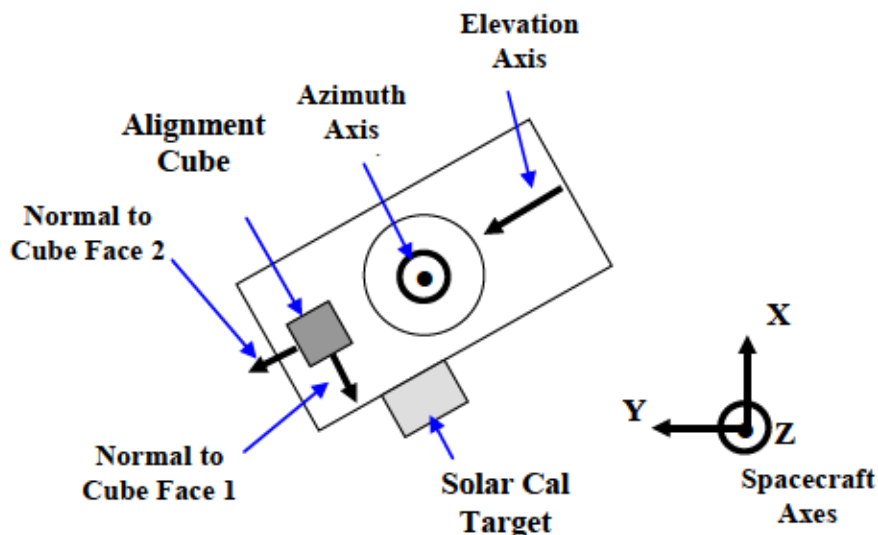


Figure 1. Schematic of DLRE instrument in its stowed position, showing relevant vectors and axes. The azimuth axis and the spacecraft Z axis point up out of the page.

Spacecraft Axis	Azimuth Rotation Axis	Elevation Rotation Axis (Stowed Position)	Normal to Cube Face 1	Normal to Cube Face 2
X	0	$-\sin(30^\circ) = -0.5$	$-\cos(30^\circ) = -0.8660$	$-\sin(30^\circ) = -0.5$
Y	0	$\cos(30^\circ) = 0.8660$	$-\sin(30^\circ) = -0.5$	$\cos(30^\circ) = 0.8660$
Z	1	0	0	0

Table 4. Nominal unit vectors relative to the spacecraft axes.

V. Laser Tracker Measurements of Rotation Axes and Solar Calibration Target

A laser tracker was used to measure the two rotation axes and the normal to the solar calibration target surface, all with respect to the alignment cube. The cube face unit normal vectors were determined as shown in Figure 2. Targets were positioned such that the laser tracker could see the target reflected in the cube face as well as directly. By determining the two vectors to each target, the reflecting cube face orientation could be determined.

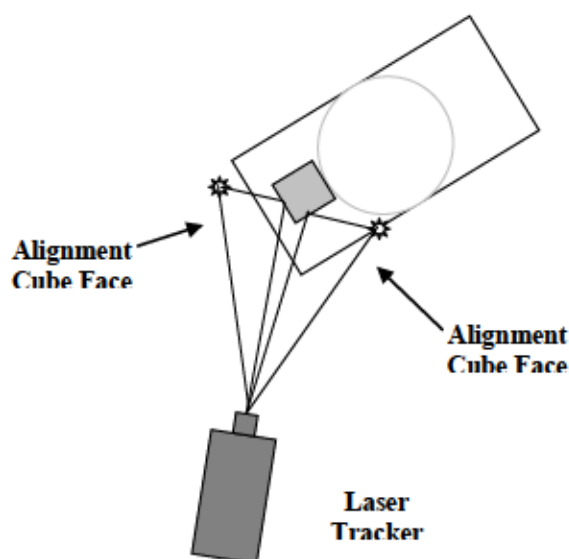


Figure 2. Laser tracker measurement of cube face normal vectors.

The rotation axes were determined by placing a target in an appropriate spot on the DLRE structure and rotating DLRE about the relevant axis. The elevation axis was measured when the azimuth axis was in stowed position (0° , azimuth actuator step 1000). The normal vector to the solar calibration target plane was determined by touching a target to each of the four corners of the solar calibration target top surface and measuring the target location in each case. The relative unit normal vectors thus determined are given in Table 5.

Axis	Azimuth Rotation Axis	Elevation Rotation Axis	Normal to Cube Face 1	Normal to Cube Face 2	Normal to Solar Calibration Target Surface
I	0.00242161	-0.00276607	-1.00000000	-0.00020120	-0.25401060
J	-0.01630531	0.99986960	0.00000146	0.99999998	-0.02245508
K	0.99986413	0.01590982	0.00000052	0.00000002	0.96694073

Table 5. Relative unit vectors determined by laser tracker measurements.

After alignment to the LRO spacecraft, we will want all alignment measurements to be described in terms of the spacecraft axes. In this preliminary analysis, we will rotate the frame of reference such that the azimuth axis is pointed along +Z and the elevation axis is pointed, to the greatest extent possible, in the X-Y plane and rotated 30° from +Y towards the -X axis. This rotation can be obtained by multiplying the original unit vectors by the rotation matrix:

$$\begin{bmatrix} 0.86740210 & -0.49750304 & -0.01021384 \\ 0.49760198 & 0.86730898 & 0.01293850 \\ 0.00242161 & -0.01630531 & 0.99986413 \end{bmatrix}$$

Multiplication by this rotation matrix results in the unit vectors given in Table 6.

Axis	Azimuth Rotation Axis	Elevation Rotation Axis	Normal to Cube Face 1	Normal to Cube Face 2	Normal to Solar Calibration Target Surface
I'	0.00000000	-0.49999996	-0.86740283	-0.49767755	-0.21903403
J'	0.00000000	0.86602533	-0.49760071	0.86720884	-0.13336091
K'	1.00000000	-0.00040222	-0.00242111	-0.01630578	0.96656037

Table 6. Laser tracker unit vectors after rotation.

VI. Theodolite Measurement From Alignment Cube to Target Projector Slit

Figure 3 shows the geometry for the theodolite measurements. Three theodolite positions were used. A Porro prism, fixed in position throughout the measurements, provided a reference angle. Table 7 lists all theodolite measurements. With the theodolite in position *A*, cube face 1 and the Porro prism were measured. With the theodolite in position *B*, cube face 2 and the Porro prism were measured. With the theodolite in position *C*, the target projector slits and the Porro prism were measured. The slit angular positions were determined by looking into the collimated beam from the target projector. A flashlight illuminated the slits to make them visible through the theodolite.

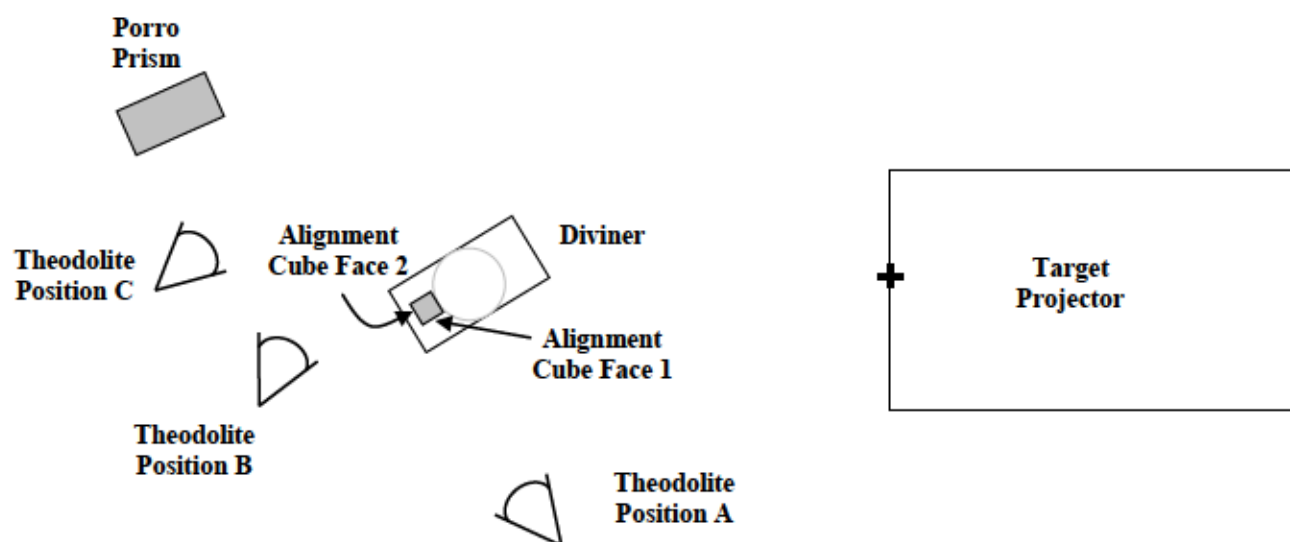


Figure 3. Schematic diagram of alignment setup, showing positions of the DLRE alignment cube, the target projector target, the Porro prism, and the three theodolite positions.

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Measurement Number	Measurement	Theodolite Position	Azimuth Angle (°)	Elevation Angle (°)
1	Cube Face 1	A	10.0001 10.0004 10.0003 10.0008 9.9999	90.2328 90.2328 90.2324 90.2336 90.2324
2	Porro Prism	A	353.9708 353.9708 353.9709 353.9707 353.9710	
3	Cube Face 2	B	10.0000 10.0020 10.0015 10.0009 10.0006	90.6231 90.6226 90.6233 90.6231 90.6232
4	Porro Prism	B	263.9701 263.9702 263.9705 263.9704 263.9698	
5	Vertical Slit in Target Projector, Instrument Removed	C	336.0800	
6	Porro Prism	C	214.7043 214.7041 214.7045 214.7042 214.7044	
7	Vertical Slit in Target Projector, Instrument Removed	C	336.0795	
8	Horizontal Slit, Reposition slit between measurements	C		88.7281 88.7160
9	Vertical Slit, Reposition slit between measurements	C	336.0450 336.0864	

Table 7. Summary of theodolite measurements.

The theodolite specified accuracy is 0.5 arcsec. Mark Thompson, the metrologist who made the measurements, feels that a more practical number for accuracy is 2 arcsec. The measurement with the most scatter (cube face 2) has a scatter of 0.002 degrees or 7 arcsec. It seems reasonable, therefore, to conclude that the accuracy of these measurements is at least better than ± 10 arcsec ($\pm 0.003^\circ$, ± 0.05 mrad).

The target projector slits are in a stepper-motor controlled wheel. It was found that the slit wheel did not always stop at the same position for a given slit. Between the various measurements of each slit, the slit was repositioned by rotating the slit wheel at least one revolution. The extremes of the slit positions were found and included in these measurements. The slit position was taken as the midpoint of these extreme measured values. Thus, from theodolite position C the slit center is $336.066 \pm 0.021^\circ$ in azimuth and $88.722 \pm 0.006^\circ$ in elevation.

Measurements 1 and 2 give an azimuth angle between cube face 1 and the Porro prism of 16.0295° (Figure 4). Measurement 1 shows that cube face 1 has an angle of 89.8600° with respect to gravity.

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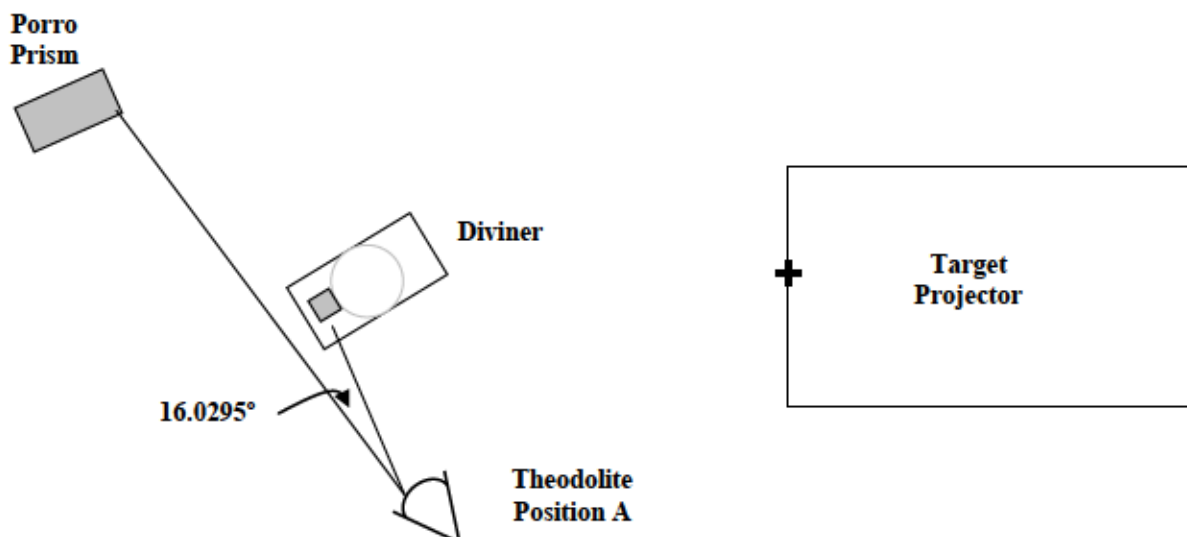


Figure 4. Relationship between cube face 1 and the Porro prism.

Measurements 3 and 4 give an azimuth angle between cube face 2 and the Porro prism of 106.0308° (Figure 5). Measurement 3 shows that cube face 2 has an angle with respect to gravity of 90.6231° .

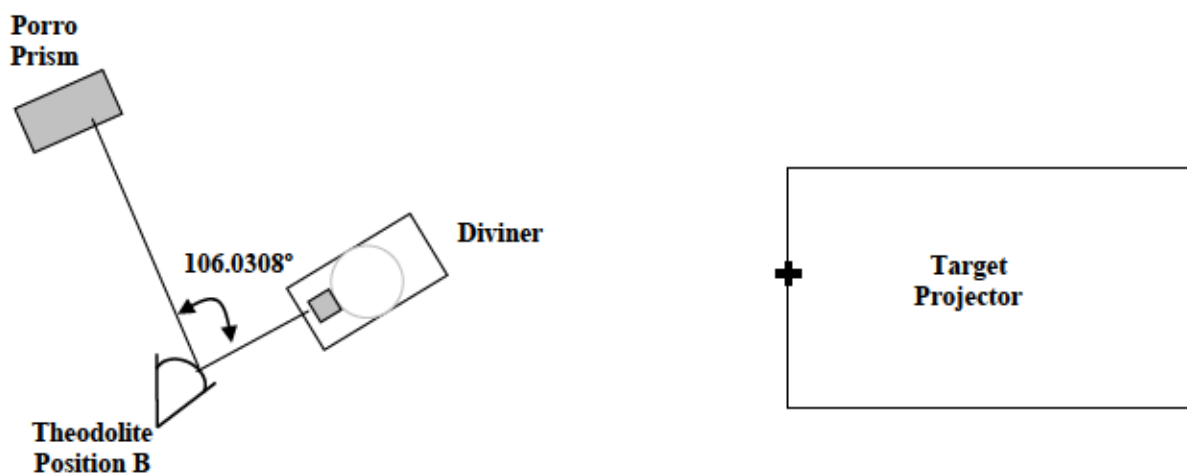


Figure 5. Relationship between cube face 2 and the Porro prism.

The measured azimuth angle between cube faces 1 and 2 is therefore $106.0308^\circ - 16.0295^\circ = 90.0013^\circ$.

Measurements 5, 6, 7, and 9 give an azimuth angle between the target projector slit and the Porro prism of 121.368° (Figure 6).

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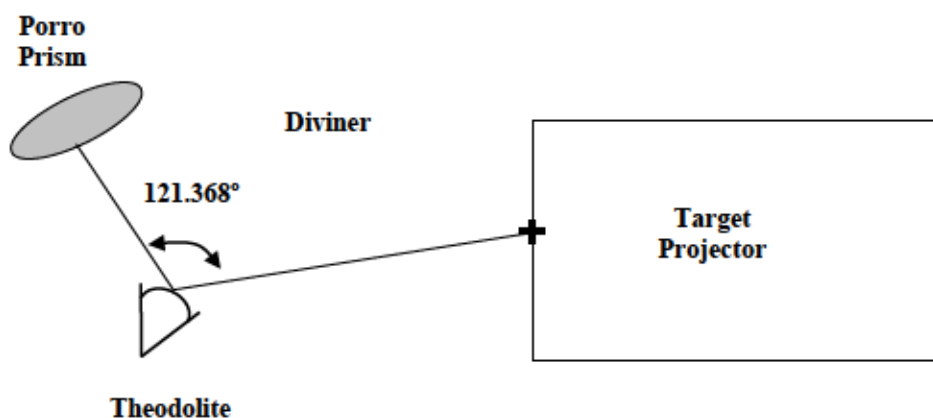


Figure 6. Relationship between the target projector slit and the Porro prism.

Measurement 8 gives an angle between the target projector slit and gravity of 88.722° . A summary of the theodolite results is given in Table 8.

Angle	Azimuth Angle ($^\circ$)	Elevation Angle ($^\circ$)
Cube Face 1 to Porro Prism	16.0295 ± 0.003	
Cube Face 2 to Porro Prism	106.0308 ± 0.003	
Target Projector Slit to Porro Prism	121.368 ± 0.021	
Cube Face 1 to Gravity		89.860 ± 0.003
Cube Face 2 to Gravity		90.623 ± 0.003
Target Projector Slit to Gravity		88.722 ± 0.006

Table 8. Summary of theodolite measurements.

We arbitrarily define a zero azimuth angle as 46° from the Porro prism, with positive angle counterclockwise in the figures. This puts the zero in azimuth roughly parallel to the spacecraft X axis. The zero in elevation angle is set to 90° from the gravity vector, with positive angles upward. The resulting angles are summarized in Table 9.

Angle	Azimuth Angle ($^\circ$)	Elevation Angle ($^\circ$)
Normal to Cube Face 1	209.9705	0.2328
Normal to Cube Face 2	119.9692	0.6231
Target Projector Slit	284.632	-1.278

Table 9. Summary of theodolite measurements.

A set of unit vectors summarizing the theodolite measurements is shown in Table 10.

Axis	Cube Face 1 Normal	Cube Face 2 Normal	Target Projector Slit
i	-0.86627557	-0.49950485	0.25254695
j	-0.49954992	0.86624283	-0.96732755
k	0.00406312	0.01087493	-0.02230346

Table 10. Unrotated unit vectors.

To put these vectors in the same frame of reference as the laser tracker measurements, they can be multiplied by the rotation matrix:

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$$\begin{bmatrix} 0.99981322 & 0.00236375 & -0.01918194 \\ -0.00197284 & 0.99979051 & 0.02037245 \\ 0.01922607 & -0.02033080 & 0.99960843 \end{bmatrix}$$

The resulting unit vectors deduced from the theodolite measurements are listed in Table 11.

Axis	Cube Face 1 Normal	Cube Face 2 Normal	Target Projector Slit
i'	-0.86737252	-0.49757257	0.25064108
j'	-0.49765347	0.86726836	-0.96807752
k'	-0.0024373030	-0.01634425	0.00222731

Table 11. Rotated unit vectors determined from theodolite measurements.

As a partial check of the measurements and axes rotations, the cube face unit vectors can be compared. The angle between two unit vectors is the arccosine of the dot product of any two vectors. The laser tracker data from Table 6 indicates that the two cube faces differ from a 90° separation by 200 μ rad. The theodolite data from Table 11 indicates that the two cube faces differ from a 90° separation by 21 μ rad. These results give some indication of the accuracy of the two types of measurements - ± 100 μ rad for the laser tracker and ± 10 μ rad for the theodolite.

After rotation, the two cube face 1 vectors, measured with the laser tracker and the theodolite, are 63 μ rad apart. After rotation the two cube face 2 vectors are 125 μ rad apart. Thus the rotations used aligned the measurement to within the laser tracker accuracy.

VII. Errors in Alignment Knowledge within the DLRE Instrument

Table 12 lists values for the primary alignment error sources within the DLRE instrument.

Error Source	Error (°)	Error (arc-sec)	Error (mrad)
Accuracy of alignment measurements (largest inconsistency is orthogonality of cube faces in laser tracker measurement).	0.006	41	0.10
Uncertainty in position of target projector slit	0.021	70	0.37
Accuracy and repeatability of azimuth actuator position (from <i>DLRE Actuator Verification Summary JPL D-41559</i>)	0.020	70	0.35
Accuracy and repeatability of elevation actuator position (from <i>DLRE Actuator Verification Summary JPL D-41559</i>)	0.020	70	0.35
Total error due to actuators (rms of previous two rows, since errors are orthogonal)	0.028	99	0.49
Error due to pixel IFOV determination (from <i>DLRE FOV Test Report JPL D-42956</i> Section X, excluding actuator portion to avoid double counting of this error)	0.040	144	0.7
Sum of errors	0.095	354	1.7
RMS of errors	0.054	194	0.94

Table 12. Errors in alignment knowledge.

VIII. Errors in Alignment Knowledge After In-Flight Calibration

During the commissioning phase of the mission, scans of the lunar horizon and comparison to LROC and LOLA data will be used as a calibration of DLRE alignment. The final DLRE alignment knowledge and accuracy will be limited by the accuracy of this calibration, of variations in DLRE's alignment with the spacecraft (spacecraft thermal distortion, jitter, etc.), and DLRE's internal variations in alignment. Careful in-flight calibration should provide better than $\pm 1/10$ pixel (± 0.6 mrad) accuracy at the time of calibration. This calibration will remove many of the sources of error listed in Table 12, such as uncertainty in the slit position during FOV measurements and some of the uncertainty in IFOV determinations. This calibration will also remove any additional FOV shifts during launch. Contributions to DLRE's internal alignment variation are listed in Table 13.

Error Source		Error (mrad)
Accuracy and repeatability of azimuth actuator position (from <i>DLRE Actuator Verification Summary JPL D-41559</i>)		0.35
Accuracy and repeatability of elevation actuator position (from <i>DLRE Actuator Verification Summary JPL D-41559</i>)		0.35
Total error due to actuators (rms of previous two rows, since errors are orthogonal)		0.49
Temperature dependence of IFOVs with respect to base of instrument (see <i>DLRE FOV Test Report, JPL D-42956, Section IX</i>).		0.53
RMS of errors		0.72

Table 13. Contributions to DLRE's internal alignment variation.

IX. Verification of Alignment Requirements

DLRE-MISC-NEW: *Diviner shall stay within its pointing accuracy allocation as stated in Table 5-1 of the LRO Pointing and Alignment Spec (431-SPEC-000113).*

Table 12 shows that the uncertainty in pointing accuracy is no worse than 1.7 mrad, compared to the 2.0 mrad shown in Table 2 (Table 5-1 of the *LRO Pointing and Alignment Spec (431-SPEC-000113)*).

DLRE-MISC-NEW: *Diviner shall stay within its pointing knowledge allocation as stated in Table 5-2 of the LRO Pointing and Alignment Spec (431-SPEC-000113).*

DLRE-MISC-NEW: *Diviner shall stay within its pointing calibration allocation as stated in Table 5-2 of the LRO Pointing and Alignment Spec (431-SPEC-000113).*

After alignment to spacecraft, Diviner's pointing knowledge (1700 μ rad per Table 12) will exceed the requirement (600 μ rad per Table 3). In-flight calibration, however, will almost bring the alignment knowledge (720 μ rad per Table 13) in line with the requirement. This smaller exceedance consumes about 1/3 of the unallocated margin in Table 3. Thus, if the spacecraft doesn't exceed its allocation significantly, the overall alignment knowledge after in-flight calibration will be within specifications.

DLRE-FRD-63a: *The orientation of the azimuth axis relative to the alignment cube shall be measured to an accuracy of ± 0.1 mrad.*

DLRE-FRD-63b: *The orientation of the elevation axis relative to the alignment cube shall be measured to an accuracy of ± 0.1 mrad.*

The axes' orientations relative to the alignment cube were measured with the laser tracker. Laser tracker measurements of the two cube faces suggest an accuracy of about ± 0.1 mrad.

DLRE-FRD-63c: *For a single orientation in azimuth and a nominally 90° orientation in elevation, alignment of the boresight of telescope A relative to the alignment cube shall be measured to better than ± 0.6 mrad.*

The field-of-view measurements represent a small range of azimuth positions and an elevation position of roughly 90° . Table 12 lists the uncertainty in pointing of a single pixel of 1.7 mrad. The boresight is determined somewhat better than this as it is an average of all pixels in each focal plane. After in-flight calibration, this error decreases to 0.72 mrad or less per Table 13. As stated above, this pointing knowledge will allow Diviner's higher-level pointing requirements to be met.

DLRE-FRD-63d: *For a single orientation in azimuth and a nominally 90° orientation in elevation, alignment of the array axes relative to the instrument axes shall be measured to better than ± 6 mrad. The axis of an array is defined by the best fit line through the peaks of all the detectors in that array.*

Figure 12 of the *DLRE FOV Test Report JPL D-42956* shows all the pixel positions in azimuth and elevation steps. This data is taken from the low-resolution field-of-view measurements. Each single measurement is accurate to 1.7 mrad (0.96 actuator steps) in each axis (Table 12). Each data point represents an average of about four single measurements, and there are 21 data points for each channel. The linear fits shown in this Figure have an uncertainty in y value (azimuth position) of about $1.7 \text{ mrad} / (4 \times 21) = 0.2 \text{ mrad}$, and an uncertainty in slope of about $0.2 \text{ mrad} / 15 = 13 \text{ } \mu\text{rad}$ (15 is roughly the rms value of elevation position about its mean). Thus the slopes are determined to much better than the required ± 6 mrad.

DLRE-FRD-240: *The plane of the solar calibration target shall be inclined 15° above the spacecraft XY plane.*

Once mounted on the spacecraft, the azimuth axis will be nearly parallel to the spacecraft Z axis (can be controlled to 6 mrad or 0.34° per Table 2). The angle of the solar calibration target above the spacecraft XY plane is the same as the angle between the solar calibration target normal and the spacecraft Z axis. This angle can be calculated by taking the arccosine of the dot product of the solar calibration target unit normal vector and the azimuth unit normal vector. Using values from Table 5 gives an angle of 14.86° .